

# A new era for the THEMIS Solar Telescope



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Methods and techniques for high-resolution spectro-polarimetry with  
current and future high-resolution telescopes:  
lessons learned and ways forward

# THEMIS Factsheet



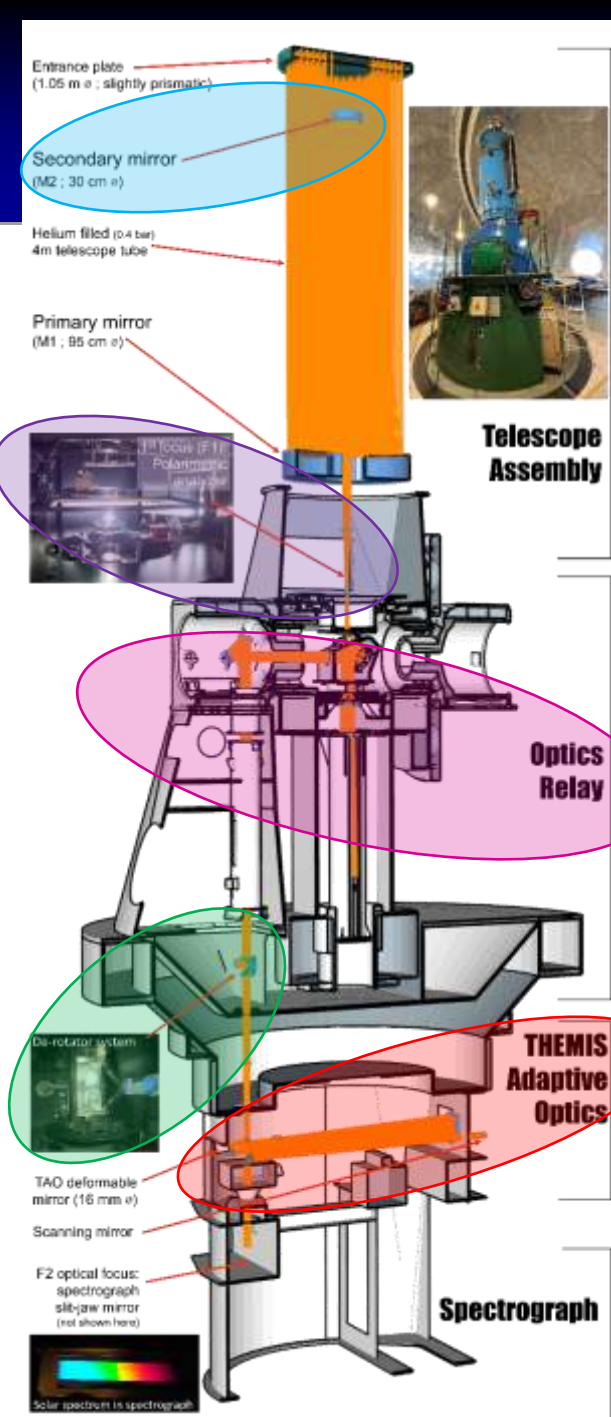
Very well-maintained but,  
be that at it may,  
a pre-AO 20<sup>th</sup> century instrument !



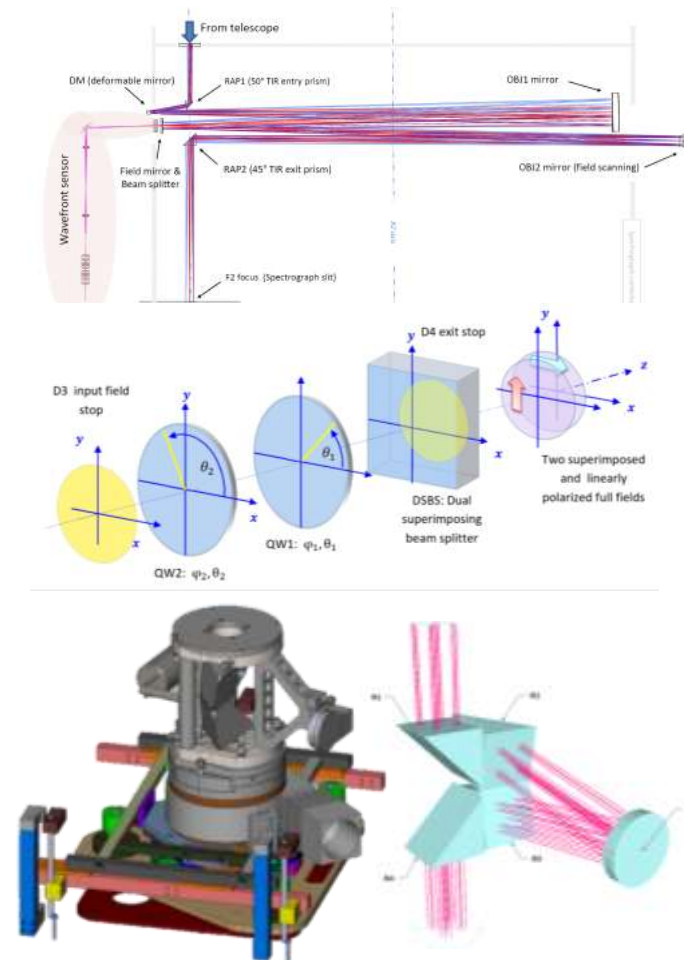
THEMIS @ OT in June 2025

- **Main french solar telescope**
  - Designed by J. Rayrole, P. Mein & M. Semel
  - Located at Teide Observatory, Tenerife, Spain
  - 1<sup>st</sup> light in March 1996, & commissioned in 1999
- 1m-class solar telescope, with one the world “slowest” optical design:
  - Aperture: 92 cm ; Effective focal length: 57m
  - Effective focal ratio: f/62
  - 60”x60” to 120”x120” square field-of-view
- **MuTi Ray spectrograph (MTR2): ideal for high-spectral resolution spectropolarimetry:**
  - Working spectral range: 4000 - 11000 Å
  - **Polarization calibration free**
  - **Ultra-high spectral resolving power:**
    - $R \sim 200\,000 - 300\,000$
  - **Simultaneous observations of user-defined set of up to 6 spectral ranges:**
    - $\sim 6-7\text{ Å}$  spec. range width with  $\sim 25\text{ mÅ}$  res.

# "Total makeover" : 2016 → 2020



- **THEMIS has been widely renewed and redesigned**
  - Thanks to EU funding: ~1M€ from 2 IAC-led SOLARNET programs
- **M2 mirror re-aluminising** (WHT & THEMIS)
- **Themis Adaptive Optics:** “classical” single-DM adaptive optics based on innovative wavefront sensing and mirror commanding concepts (AIRI@CRAL & THEMIS)
- **Superimposed dual-beam polarimetric analysis** without field limitation (Semel M., Lopez A., Le Men C. & THEMIS)
- **“Polarization- friendly” complete redesign of the whole transfer optics (M3, M4 & M5)** (Le Men, C. & THEMIS)
- **New de-rotator system** (THEMIS)
- + new context, broadband and spectral cameras.

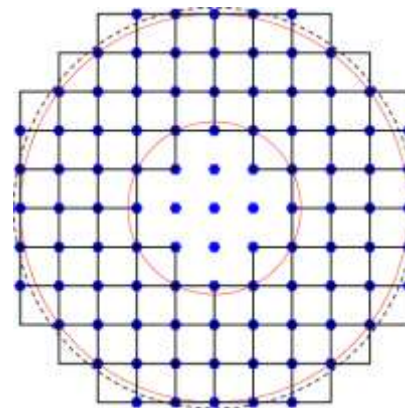




# THEMIS Adaptive Optics (TAO)



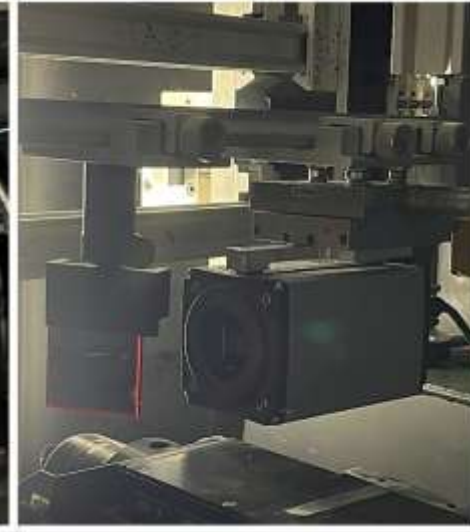
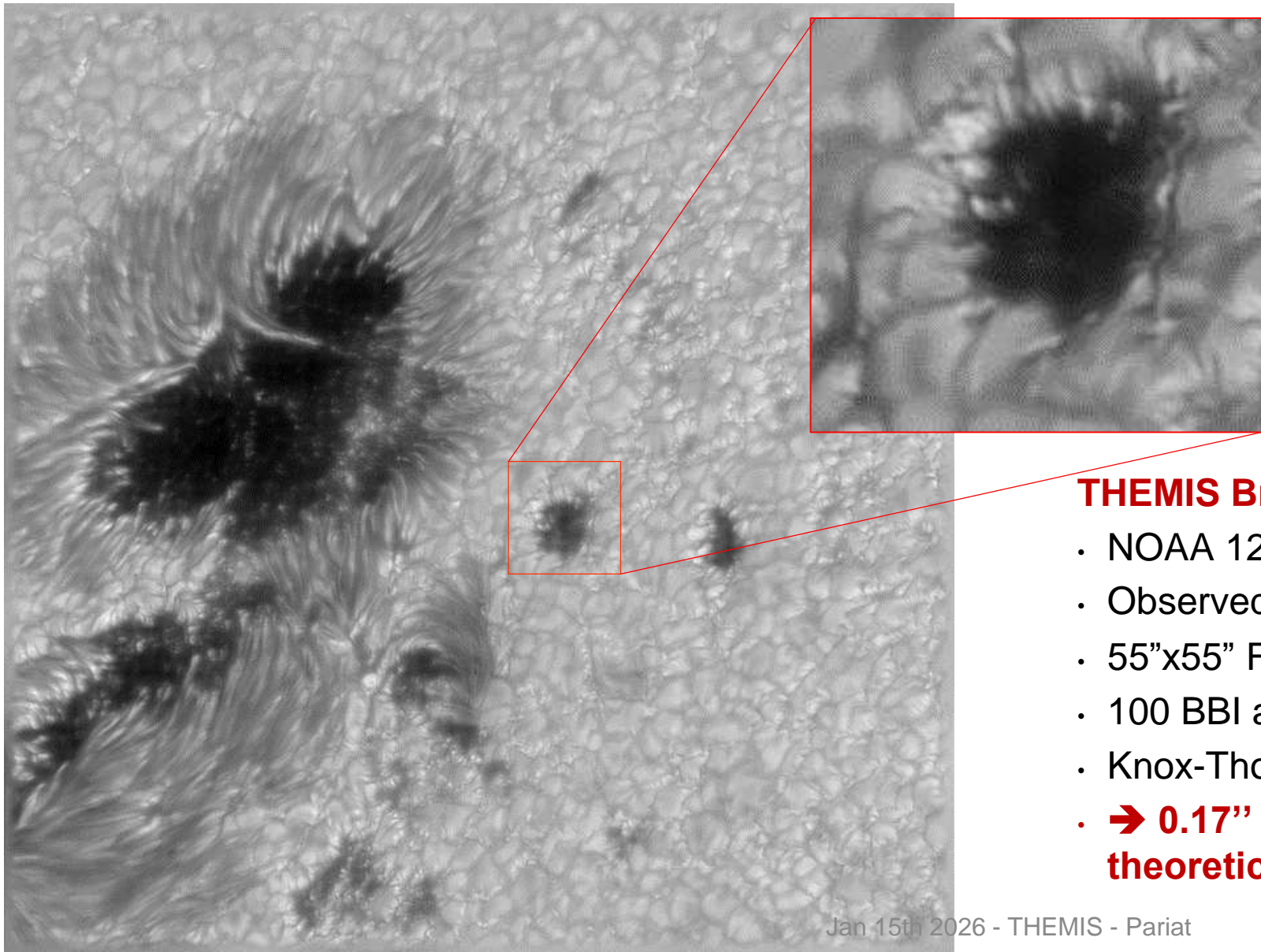
- Specifications
  - **76 sub-aperture Shack-Hartmann WFS**
    - 380×380 pixel WFS images, Mikatron EoSens 4CXP detector
  - **THEMIS-optical-path-compatible 16mm DM**
  - **97 actuators** on ALPAO deformable mirror
  - Real time correction (RTC)
    - Computer: CPU i7-4790K (Q2'14) at 4.2 GHz, 4 cores, up to 50 Gflops/core with AVX2 + FMA instructions.
- Objectives:
  - **✓ Closed AO loop on the Sun**
    - started from scratch mid-2016 → Dec. 2020
  - **✓ RTC software running in CPU @1250 Hz**
  - **Ongoing (→ winter 2025-2026): unsupervised AO system**
    - optimal correction whatever the conditions
    - provide inferred seeing conditions



TAO geometry with a combination of DM/wavefront sensor set up in 'Fried configuration' with a spacing number of 10



# THEMIS at diffraction limit: NOAA 12975

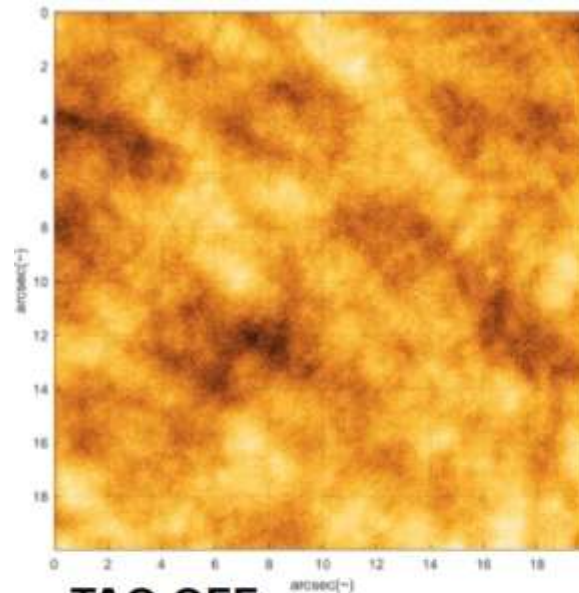
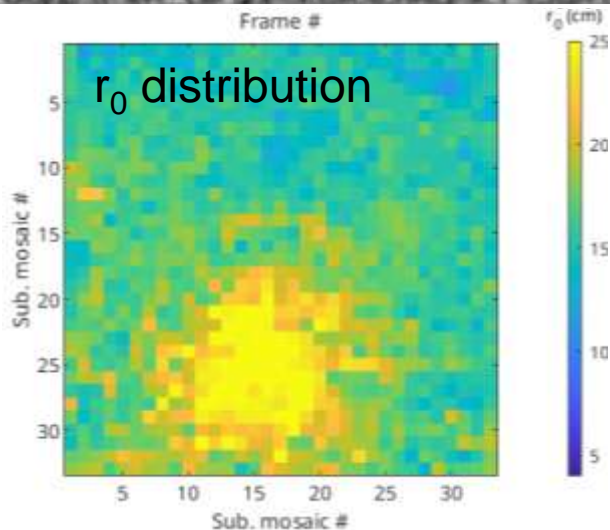
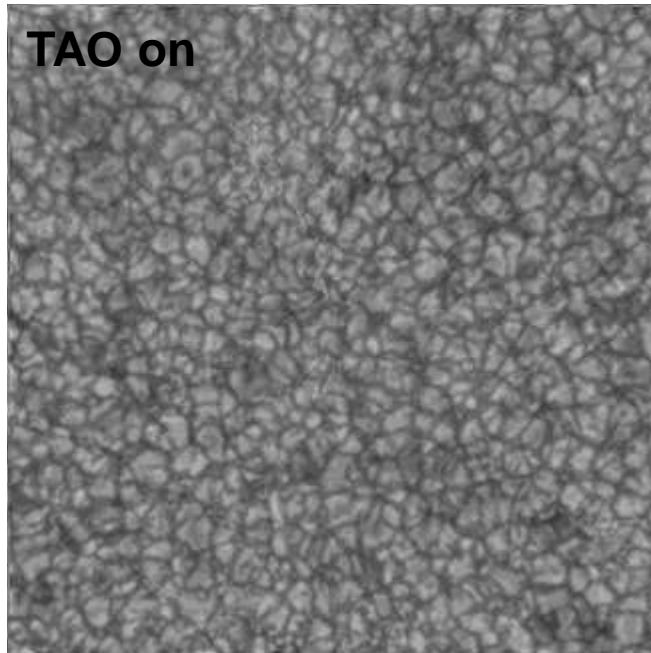


## THEMIS Broadband Imaging (BBI)

- NOAA 12975 on 2022/03/31
- Observed @  $\sim 630\text{nm}$  ; 1nm broadband red filter
- 55"x55" FOV
- 100 BBI acquisition @ 40 images/s
- Knox-Thompson (speckle) image post processing
- ➔ **0.17" resolution (0.035"/pixel) near THEMIS theoretical diffraction limit of 0.15"**

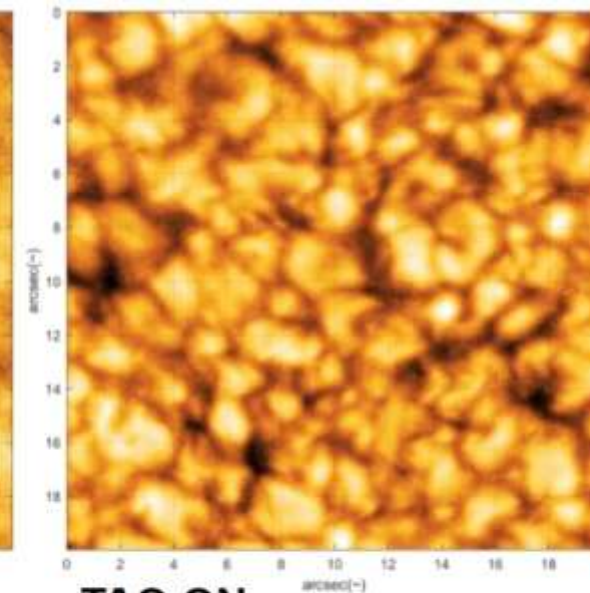


# THEMIS Adaptive Optics (TAO): results on granulation



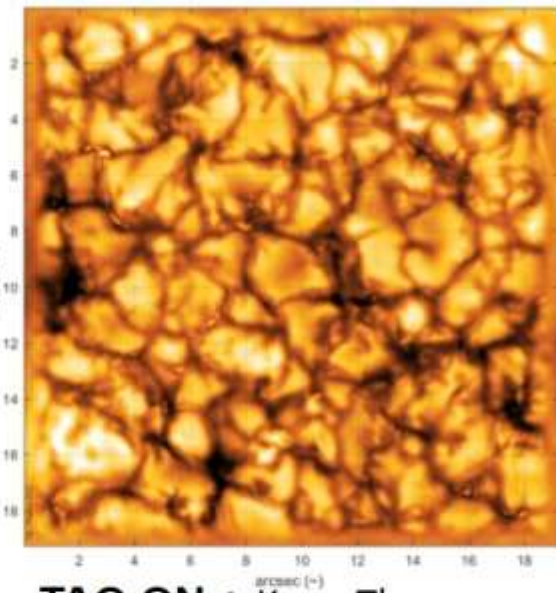
**TAO OFF**

- seeing "daytime bad" :  $r_0 \approx 3-4$  cm
- granulation contrast: 1.6 %



**TAO ON**

- seeing "daytime bad" :  $r_0 \approx 3-4$  cm
- granulation contrast: 4.2 %

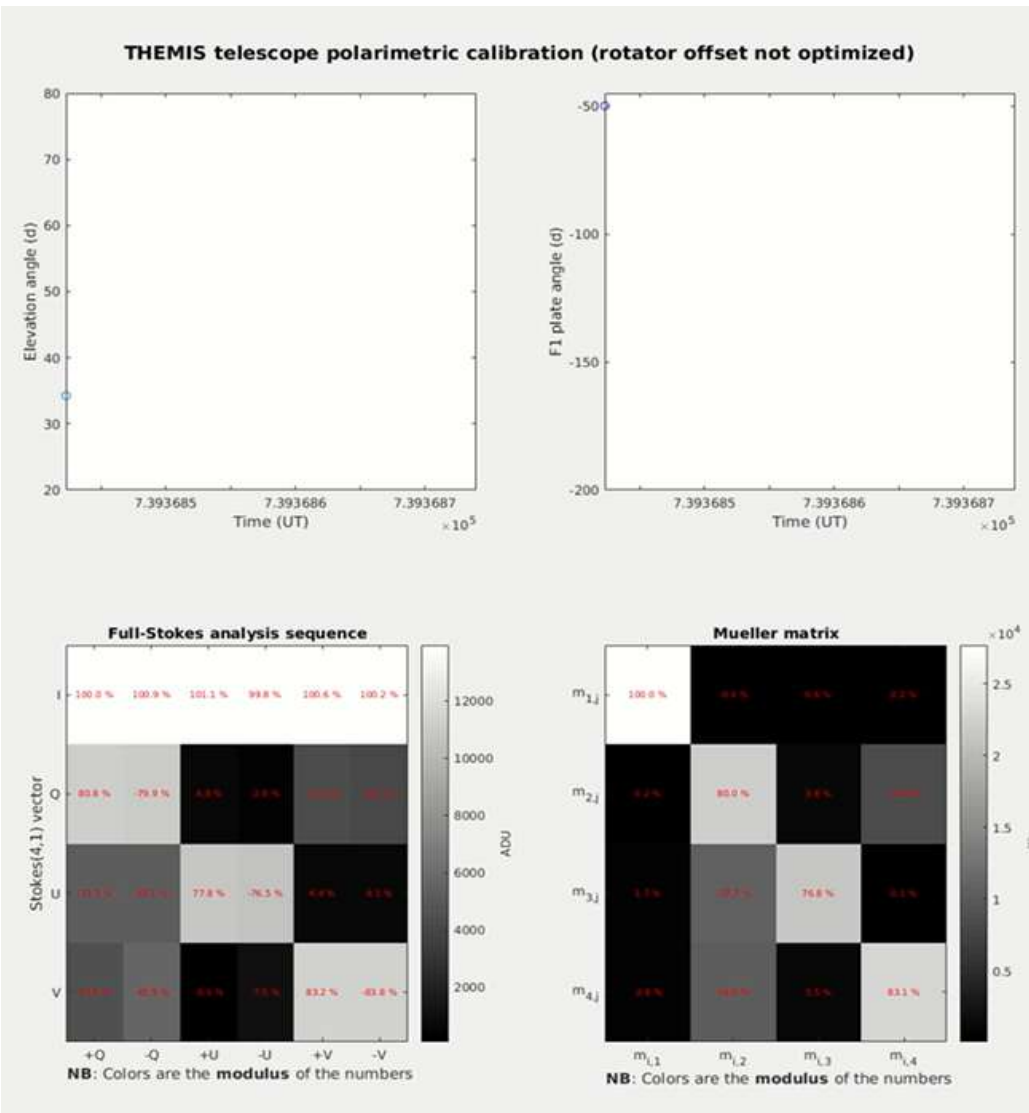


**TAO ON + Knox-Thompson  
reconstruction (100 frames)**

- granulation contrast: 9.6 %

- **TAO permits significant quantitative image quality gain:**
  - **in effective seeing:** Fried's coherence length from  $r_0 \sim 7$  cm (ave. seeing)
    - $\sim 25$  cm at TAO focus
    - $\sim 17$  cm on rest of FOV, away from isoplanatic patch
  - **in granulation contrast:** from  $\sim 1-2\%$  (bad seeing)
    - to  $\sim 10\%$  (with image reconstruction)

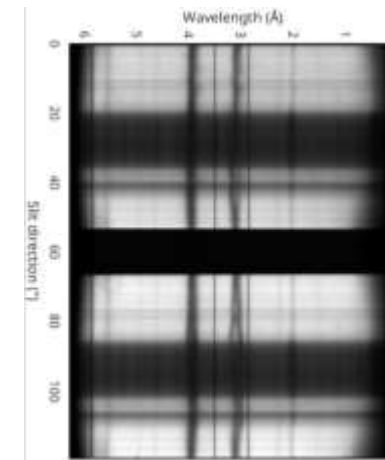
# THEMIS polarization analysis



## • New “polarization friendly”

### AO-compatible optical path:

- Polarization units + double Savart plates @ F1  
→ **dual-beam with beam exchange.**
- Wollaston prisms in front of spectral cameras:  
→ complementary Stokes on camera FOV



## • THEMIS Mueller matrix:

$$M_{THEMIS} = \begin{pmatrix} 1.000 & -0.009 & -0.003 & 0.001 \\ -0.008 & 0.885 & 0.016 & -0.033 \\ 0.014 & -0.436 & 0.872 & 0.033 \\ -0.019 & 0.415 & 0.008 & 0.873 \end{pmatrix}$$

- Averaged over one full day
- Includes changing elevation axis and field derotation
- Quite constant along one day
- **THEMIS remains a strongly polarization-calibration-free telescope, ideal for excellent spectropolarimetric measurements.**

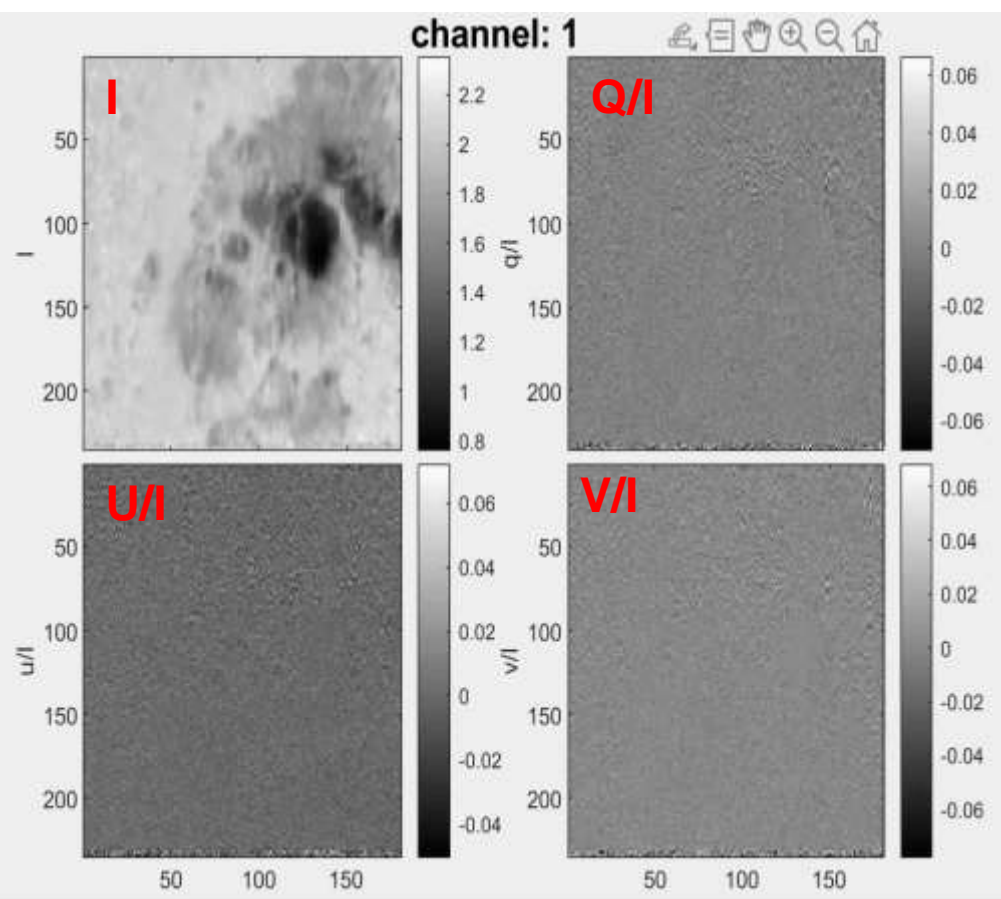


# Stokes parameters maps

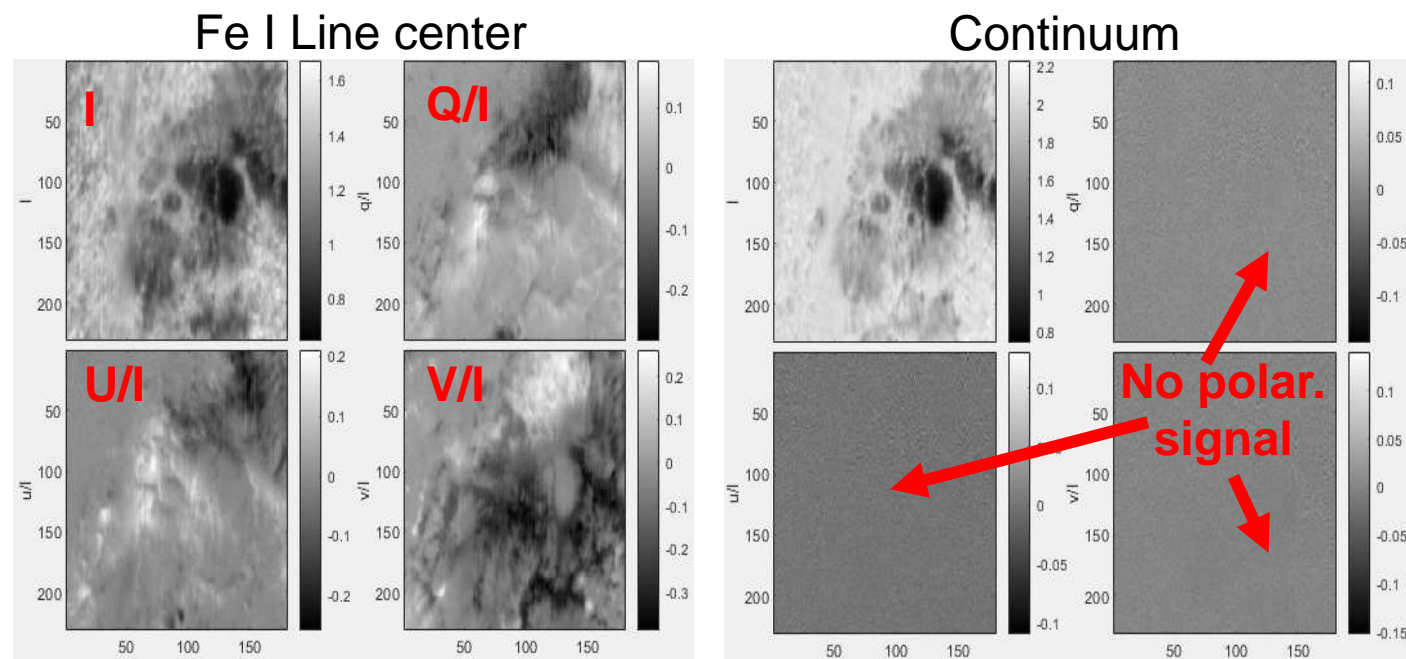
Stokes map datacube

- 256  $\lambda$  channels after 2x rebinning
- 25mÅ /pixel spectral resolution

- **Complete polarization signal now routinely measured**
  - 4D data array of 4 Stokes parameter (x, y,  $\lambda$ , S).
  - User-friendly software in development.
- **THEMIS goals : B maps with spatial resolution < 0.5''**
  - x10 area resolution improvement



NOAA 14100 ; 29/05/2025 10:30-10:50 UT



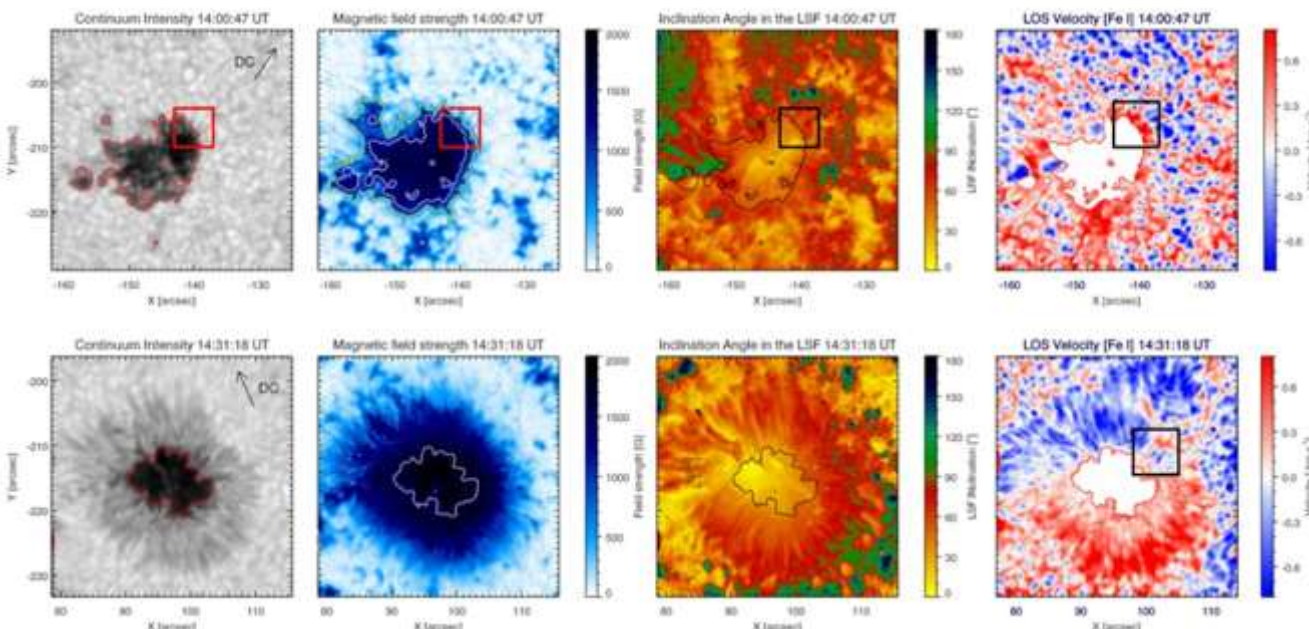


# Upcoming : IBIS 2.0 @ THEMIS



## IBIS : Interferometric Bldimensional Spectrometer

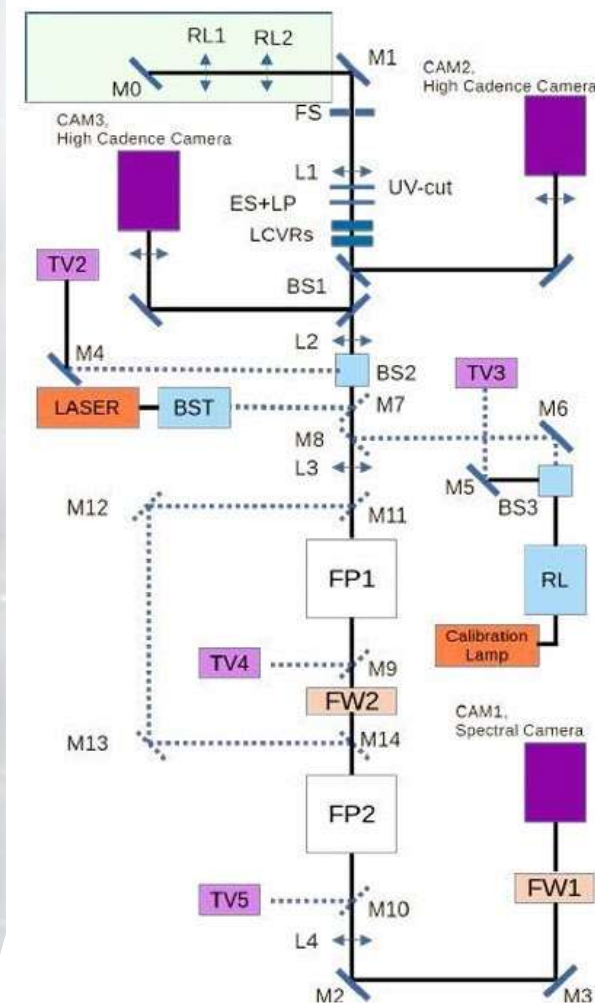
- spectro-imager ( $x, y, \lambda$ ) : dual Fabry-Perot & interference filters
- 200 000 spectral resolution
- short exposure times / polarimetric mode



Intensity, magnetic field strength, field inclination angle, and LOS velocities on 2012 May 28 (14:00-14:30 UT): before (top) and after (bottom) penumbra formation. SIR inversion of the Stokes profiles of the Fe I 630.25 nm line acquired by IBIS . (from Murabito et al. 2016)



<https://www.ibis20.inaf.it>



Ermolli et al. 2024

# Upcoming : IBIS 2.0 to THEMIS



- 2003-2019: running at the Dunn solar tower (DST) ~100 papers based on IBIS over 15 years
  - Upgraded IBIS looking for suitable telescope since 2019
  - TAO performance attractive for IBIS
  - THEMIS has no equivalent instrumental mode
  - ➔ Memorandum of Understanding signed between INAF and CNRS in winter 2024-2025
  - **Nov. 25: preparatory installation of optical bench**
  - **Spring 2026 : IBIS 2.0 installation & commissioning.**
- **IBIS2.0 is an outstanding synergic complement of THEMIS long slit spectrograph**
  - **Foster and renew French-Italian scientific collaboration** in high-resolution solar physics, beneficial at large for EU solar physics (e.g. EST)



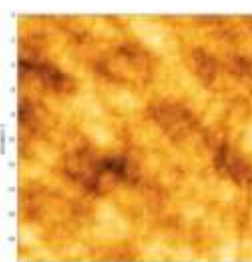
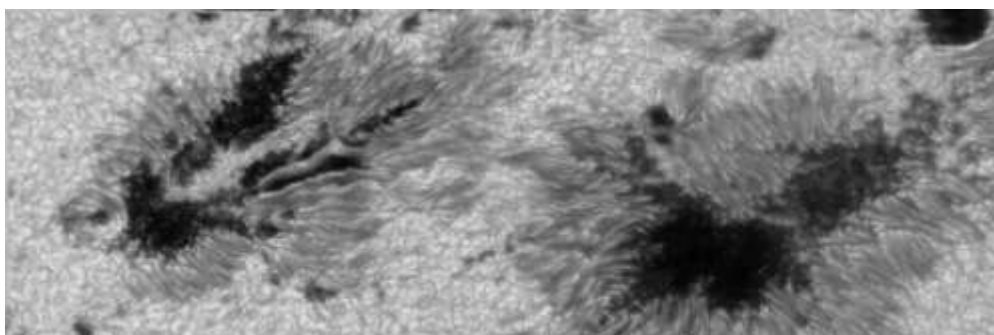
Nov. 2025



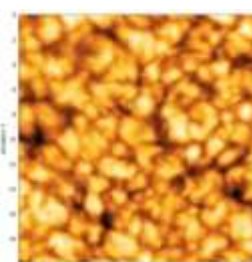
# Takeaways



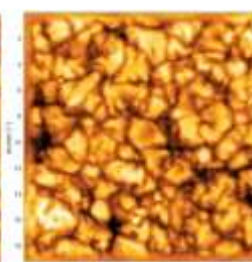
- **THEMIS is now a competitive 21<sup>st</sup> century telescope in the 1m-1.5m class with renewed spectropolarimetric capacities synergic with existing infrastructure**



**TAO OFF**  
• seeing "daytime bad" :  $\theta_s \approx 3-4$  cm  
• granulation contrast: 1.6 %

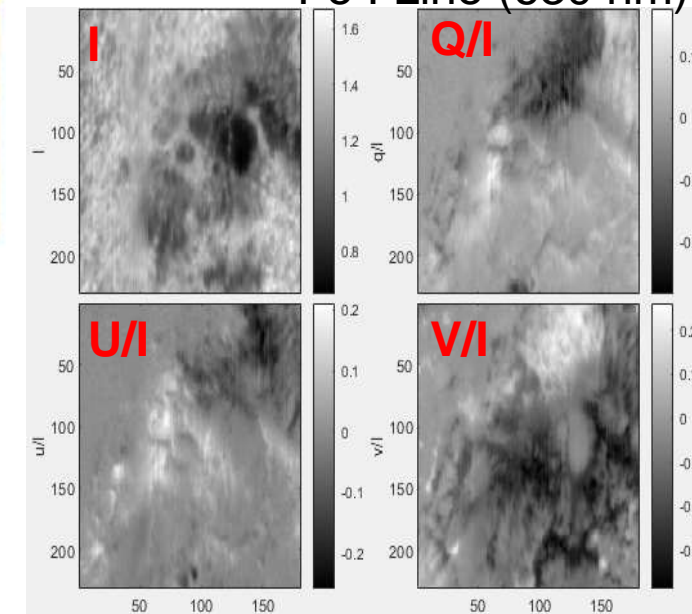


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**TAO ON + Knox-Thompson reconstruction (100 frames)**  
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Fe I Line (630 nm)



- **Installation of the IBIS 2.0 spectro-imager (spring 2026)**
- **Funded science projects**
  - hiring of early career researchers (1 PhD, 2 Postdocs) + 1 operator
  - shall strengthen French high res. solar phys. community
- **New outreach tools**
  - THEMIS website; Instagram account
- **Call for (reduced) 2026 campaign now open !**



<https://www.themis.iac.es/>